

# DYNAMICO

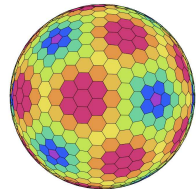
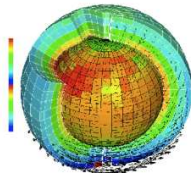
## Dynamical core on Icosahedral grid

T. Dubos<sup>146</sup>, Y. Meurdesoif<sup>356</sup>, S. Dubey<sup>2</sup>, F. Hourdin<sup>16</sup>...

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<sup>3</sup>Laboratoire des Sciences du Climat et de l'Environnement

<sup>4</sup>École Polytechnique, <sup>5</sup>CEA, <sup>6</sup>Institut Pierre Simon Laplace,



## DYNAMICO fact sheet

- hydrostatic, shallow-atmosphere
- icosahedral, hexagonal, C-grid, structured
- pressure-based hybrid terrain-following  $\eta$  coordinate
- Lorentz vertical staggering
- mass- and enstrophy-conserving FD  
(Sadourny, 1975 ; Ringler et al., 2010)
- explicit 4-th order dissipation
- Eulerian positive definite, slope-limited transport (Dubey et al., submitted)

## Planned 2012-2013

- quasi-hydrostatic, deep-atmosphere (M. Tort, PhD)
- energy conserving option
- coupling with LMD-Z physics package
- aquaplanet experiments

## 1 LMD-Z

- The LMD-Z core
- Conservation of enstrophy vs energy

## 2 DYNAMICO

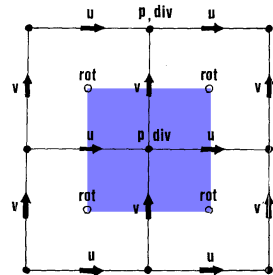
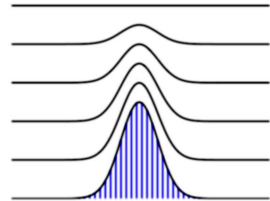
- The DYNAMICO project
- The DYNAMICO core
- First runs at DCMIP 2012

## 3 Ongoing work

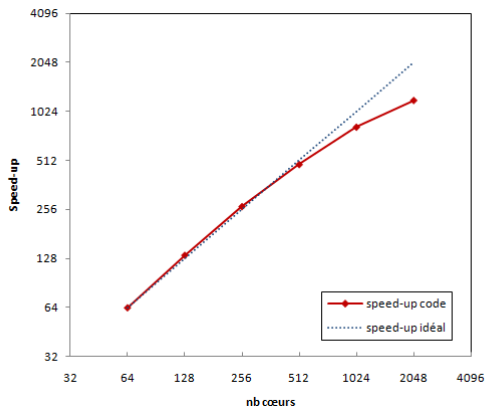
- Conservative regridding
- Deep-atmosphere dynamics

# The LMD-Z core

- hydrostatic, shallow-atmosphere
- lat-lon, C-grid + polar filters
- grid-stretching
- pressure-based hybrid terrain-following  $\eta$  coordinate
- Lorentz vertical staggering
- mass- and enstrophy-conserving (Sadourny, 1975)
- explicit 4-th order dissipation
- Eulerian positive definite, slope-limited transport (Hourdin & Armengaud, 1999)
- used to model planetary atmospheres (Mars, Venus, Titan, ...)



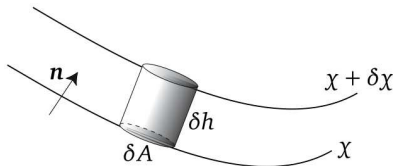
# Scalability



*Y. Meurdesoif (2010, 1/4 degree)*

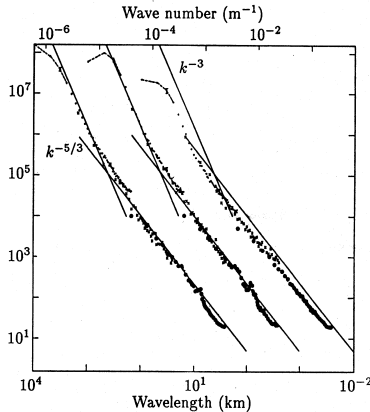
# Potential vorticity and (potential) enstrophy

- Conservation of potential vorticity implies limits on the generation of vorticity

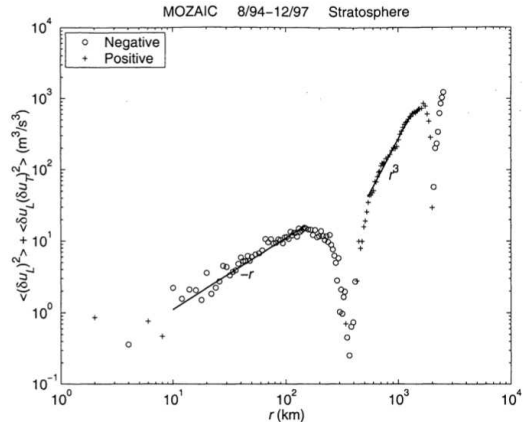


- At the discrete level, 3 levels of “vorticity conservation”
  - 1 Pressure gradient generates no vorticity
  - 2 Potential vorticity obeys an implied transport equation
  - 3 Potential enstrophy is conserved
- Conservation of energy and enstrophy tend to conflict with each other (Arakawa, 1966 ; Sadourny, 1975; Arakawa & Lamb, 1982 ; Ringler et al., 2010)

# Enstrophy vs energy



*Nastrom & Gage, 1985*



*Cho & Lindborg, 2001*

# The DYNAMICO project

## Goals & principles

- Revive an interest in numerical methods at LMD/IPSL
- Break the scalability bottleneck by moving LMD-Z to a quasi-uniform-grid
- Hydrostatic core an important milestone suitable for short-term application to climate modelling
- Provide at least the properties already present in LMD-Z
- Extend LMD-Z to deep atmospheres
- Prefer simplicity & not reinvent the wheel !

## Brief history

- 2009 : started as Indian-French project
- 2010 : work on 2D transport scheme (S. Dubey)
- 2011 : shallow-water model (Ringler et al. , 2010)
- mid-2012 : dry 3D core (Y. Meurdesoif)



# Old parts and new parts

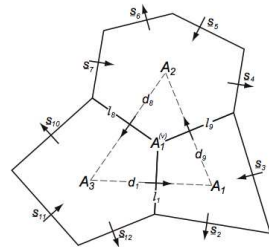
$$\frac{\partial m}{\partial t} + \frac{\partial W}{\partial \eta} + \nabla_{\eta} \cdot (\overline{m}^h u) \quad m = -\frac{1}{g} \frac{\partial p}{\partial \eta}$$

$$\frac{\partial m q}{\partial t} + \frac{\partial}{\partial \eta} (W \overline{q}^v) + \nabla_{\eta} \cdot (U \overline{q}^h) = S_q$$

$$\frac{\partial \Phi}{\partial \eta} + g \frac{m}{\overline{\rho}^v} = 0$$

$$\frac{\partial u}{\partial t} + \frac{\frac{\partial u}{\partial \eta}^v \overline{W}^{vh}}{\overline{m}^h} + \frac{(f + \nabla_{\eta} \times u) \times u}{\overline{m}^h}^{TRISK}$$

$$+ \nabla_{\eta} \left( \frac{\overline{u}^{2h}}{2} + e + \frac{p}{\rho} - \theta \pi + \Phi \right) \\ + \overline{\theta}^h \nabla_{\eta} \pi = S_u$$



*Thuburn et al., 2009*

*Miura (2007) ;*

*Dubey et al., submitted*

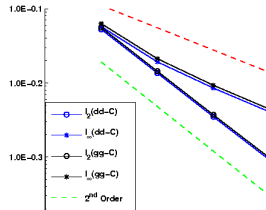
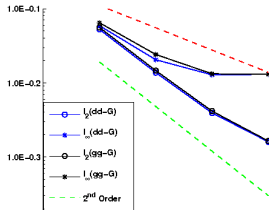
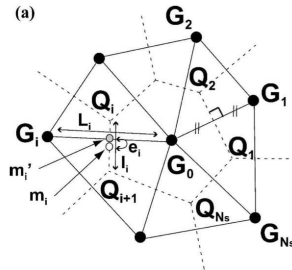
*Ringler et al., 2010 ; Dubos,  
PDEs on the Sphere 2012*

# Gradient reconstruction for slope-limited finite-volume transport

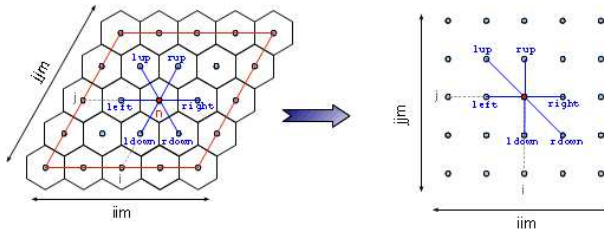
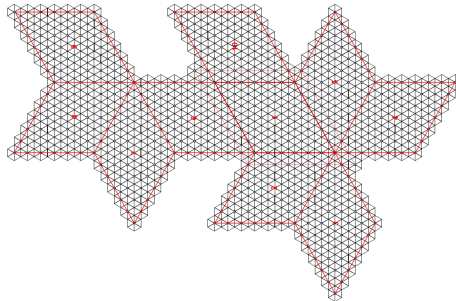
- Problem : first-order estimate of gradient given values around a given cell
- Explicit solution : Green-Gauss theorem
- Requires second-order accuracy estimate of point values

⇒ must use centroids of control volumes

Miura & Kimoto, 2005

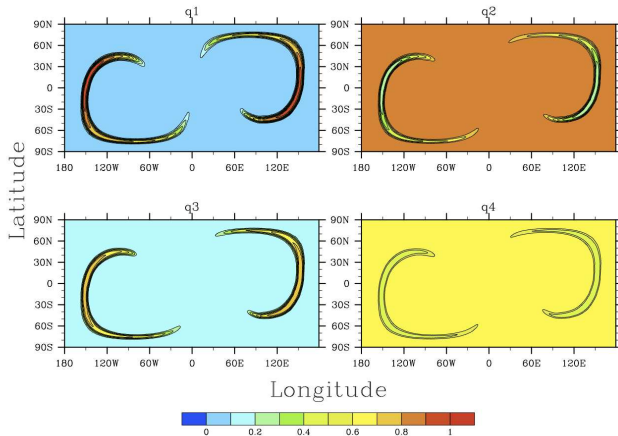


# The icosahedral grid is structured



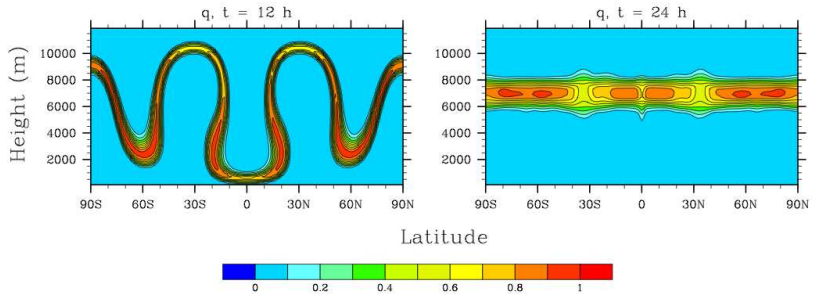
# 3D transport

DYNAMICO 4900 m,  $t = 6$  days



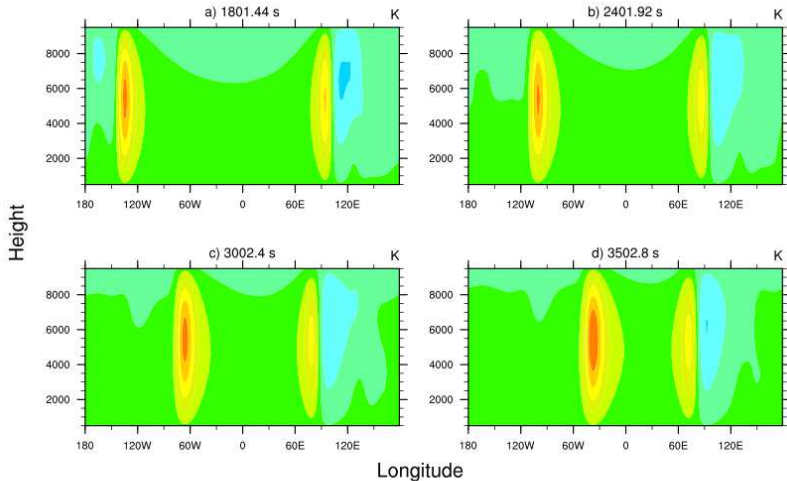
# 3D transport

DYNAMICO 1x1L60



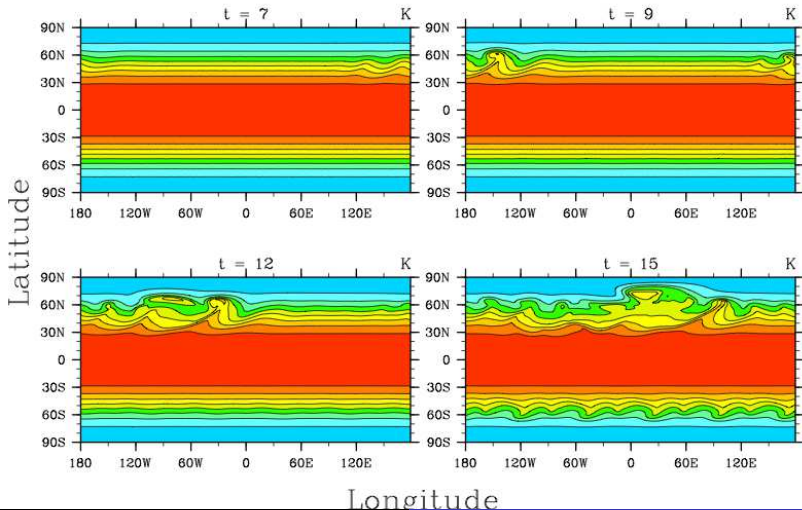
# Gravity wave

## DYNAMICO, Test 31 theta', Rotated Grid

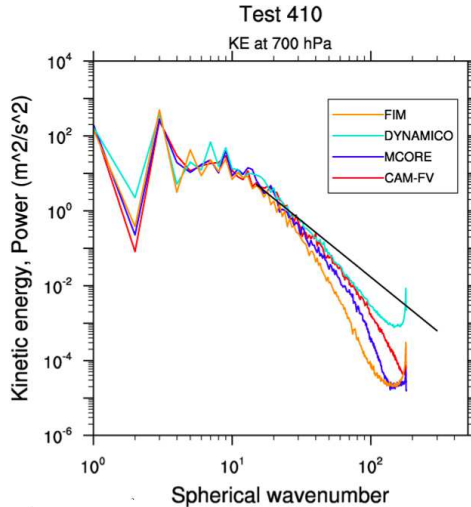


# Baroclinic wave (Jablonowski & Williamson)

DYNAMICO Test 410, theta 850 hPa



# Baroclinic wave (Jablonowski & Williamson)



(C. Jablonowski, *PDEs on the Sphere* 2012)



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- Conservation of enstrophy vs energy

## 2 DYNAMICO

- The DYNAMICO project
- The DYNAMICO core
- First runs at DCMIP 2012

## 3 Ongoing work

- Conservative regridding
- Deep-atmosphere dynamics

# Ongoing work

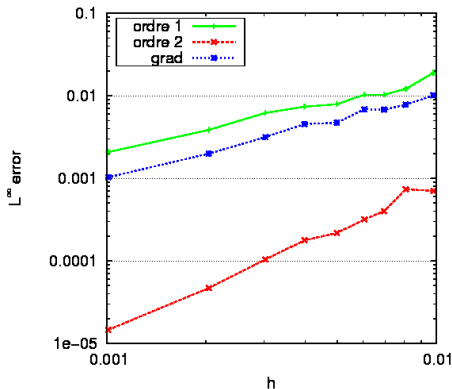
## Planned features

- Parallel I/O (XIOS, Y. Meurdesoif)
- On-the-fly parallel post-processing (with J. Thuburn, U. Exeter) starting with conservative regridding (E. Kritsikis)
- Deep-atmosphere dynamics (M. Tort)
- Grid stretching

## Potentially desirable features

- Non-orthogonal C-grid
- Conservative grid nesting (N. Kevlahan, M. Aechtner U. McMaster)
- Other approaches :
  - well-balanced finite volumes (F. Bouchut)
  - geometric schemes (F. Gay-Balmaz)
  - mixed finite elements (S. Christiansen, U. Oslo)

# Second-order conservative regridding



# Deep-atmosphere dynamics

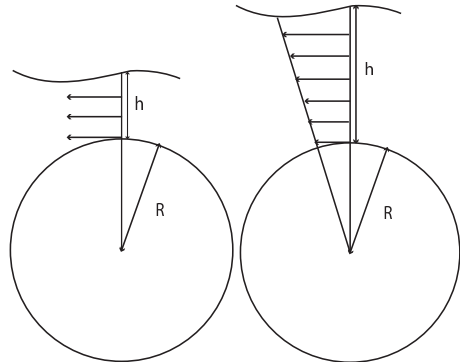
## Deep quasi-hydrostatic equations in a general vertical coordinate

- have time-dependent metric terms
- and a full Coriolis force

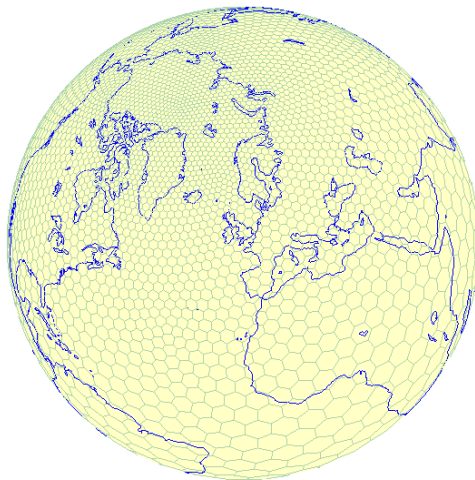
### PV/Energy Conserving formulation

Incorporate metric and entrainment velocity into prognostic variable for velocity

⇒ vector-invariant form



# Grid stretching (Schmidt transform)



# Summary

- DYNAMICO is now a (prototype) icosahedral-hexagonal hydrostatic core
- Low-order approach based on
  - variational principles  $\Rightarrow$  discrete conservation,
  - simplicity
  - reuse of suitable existing parts from IPSL or elsewhere
- Much validation / optimization still needed
- Goal is to put it to effective use as soon as possible on Earth and planets
- We expect to also improve the legacy lat-lon LMD-Z which will not disappear any time soon