




1 seismolab: A Python package for analyzing 2 space-based observations of variable stars

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6 Summary

7 Classical pulsating stars are characterized by periodic or multi-periodic brightness variations
8 from several hundredths to a few tenths of relative magnitudes. The observation of these
9 variable stars is essential for testing pulsation and stellar evolution models. The different forms
10 of frequency spectra are powerful tool for comparing observations and theoretical models.
11 Long-term period changes can reveal information about the motion of the variable star in
12 a binary system ([Plachy et al., 2021](#)). Proper classification of variable stars, which usually
13 requires rigorous analysis, is crucial to studying the properties of clear samples ([Tarczay-Nehéz
14 et al., 2023](#)).

15 Upon entering the era of photometric space missions, the launch of the NASA *Kepler* and *TESS*
16 missions have brought a new opportunity to expand the science of variable star, enabling the
17 characterization of short-term variability and the detection of millimagnitude-level variations.
18 The latest generation of large photometric and astrometric surveys has greatly expanded the
19 number of known, observed and classified variable stars. Mining these large data sets has led
20 to new discoveries and more detailed analysis in this field [Benkó et al. \(2023\)](#).

21 Several techniques have been developed to search for periodicities of the light curves and also
22 for any deviation from the strictly periodic behavior. We have developed a Python package,
23 *seismolab*, which implements various methods for downloading, analyzing, and visualizing
24 data of variable stars from space-based surveys. The framework is primary intended to be
25 used with data obtained by the *Kepler*, *TESS* and *Gaia* surveys, but can also be used by other
26 similar existing and future surveys. Some modules are also useful for analyzing ground-based
27 observations.

28 Statement of need

29 The purpose of *seismolab* is sixfold. It is able to combine *Gaia* data with the Bailer-Jones
30 distance catalog ([Bailer-Jones et al., 2021](#)), galactic extinction maps ([Bovy et al., 2016](#)) and
31 magnitudes from the Simbad catalog ([Wenger et al., 2000](#)). Different modules are implemented
32 to extract the Fourier coefficients and Fourier parameters of light curves, and to derive the
33 temporal variation of the amplitude, phase and zero point of the dominant variation ([Benkó
34 et al., 2023](#)). One can extract the minimum or maximum times and derive an O-C diagram
35 ([Sterken, 2005](#)). The package can be use to fill gaps in time series data using the method
36 of inpainting [Sandrine Pires et al. \(2015\)](#). There is also a module that provides different
37 transformation methods for time-frequency analysis ([Kolláth & Buchler, 1997](#)).

38 The documentation of *seismolab* consists of pages describing the various available functions,
39 as well as tutorial notebooks.

40 Acknowledgements

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