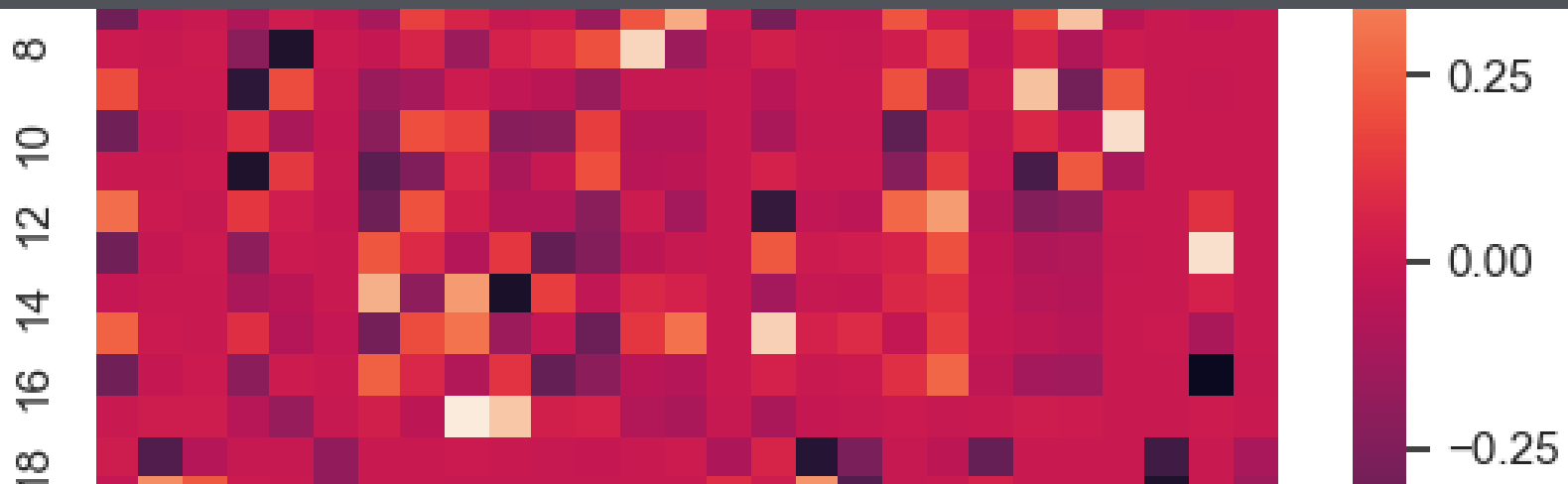


L1C ENSEMBLE FCDR



Progress update

HARMONISE → ENSEMBLE

- 9 harmonised AVHRR Easy FCDRs (MetOp-A → NOAA-11)
- 3 IR channels (3.7 μm , 11 & 12 μm)
- 3 (or 4) harmonisation coefficients (without WV) per channel per sensor → 27 (or 36) coefficients per channel
- We know the uncertainty on each coefficient
- We know their correlation matrices [27 x 27] or [36 x 36]

‘best-case’
harmonised
AVHRR Easy
FCDR

Measurement Equation for Ch3B (3.7 μm)

$$L = a_0 + ((L_{\text{ict}} * (0.985140 + a_1)) / (C_{\text{ict}} - C_s)) * (C_e - C_s) + a_2 * T_{\text{inst}} + a_3 * f(\text{WV})$$

$$L = a_0 + ((L_{\text{ict}} * (0.985140 + a_1)) / (C_{\text{ict}} - C_s) + a_2 * (C_e - C_{\text{ict}})) * (C_e - C_s) + a_3 * T_{\text{inst}} + a_4 * f(\text{WV})$$

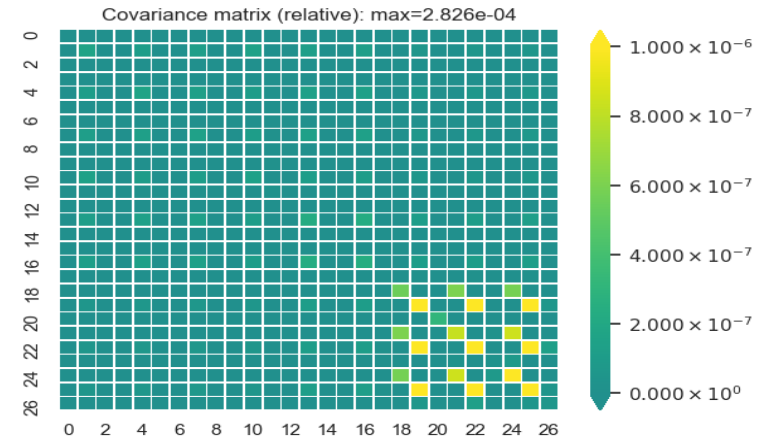
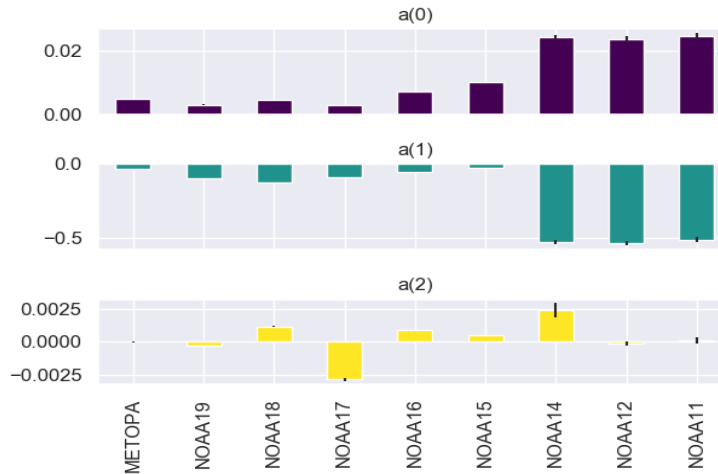
Measurement Equation for Ch4 (11 μm) & Ch5 (12 μm)

Q1: Can we generate another 10 FCDRs within the range of uncertainty but having a similar inter-sensor correlation structure to the best-case?

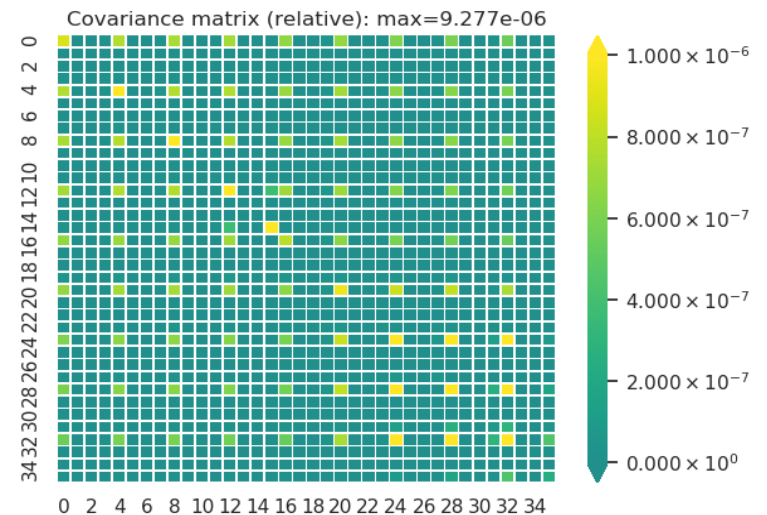
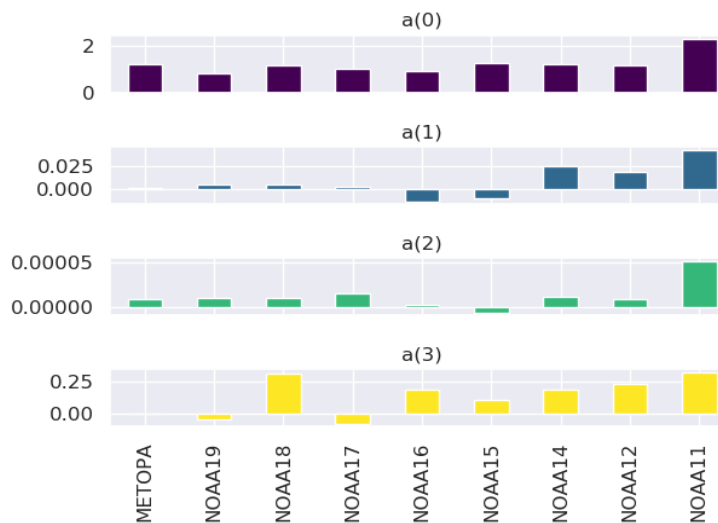
Q2: Why bother?

THE BEST CASE

3.7 μm



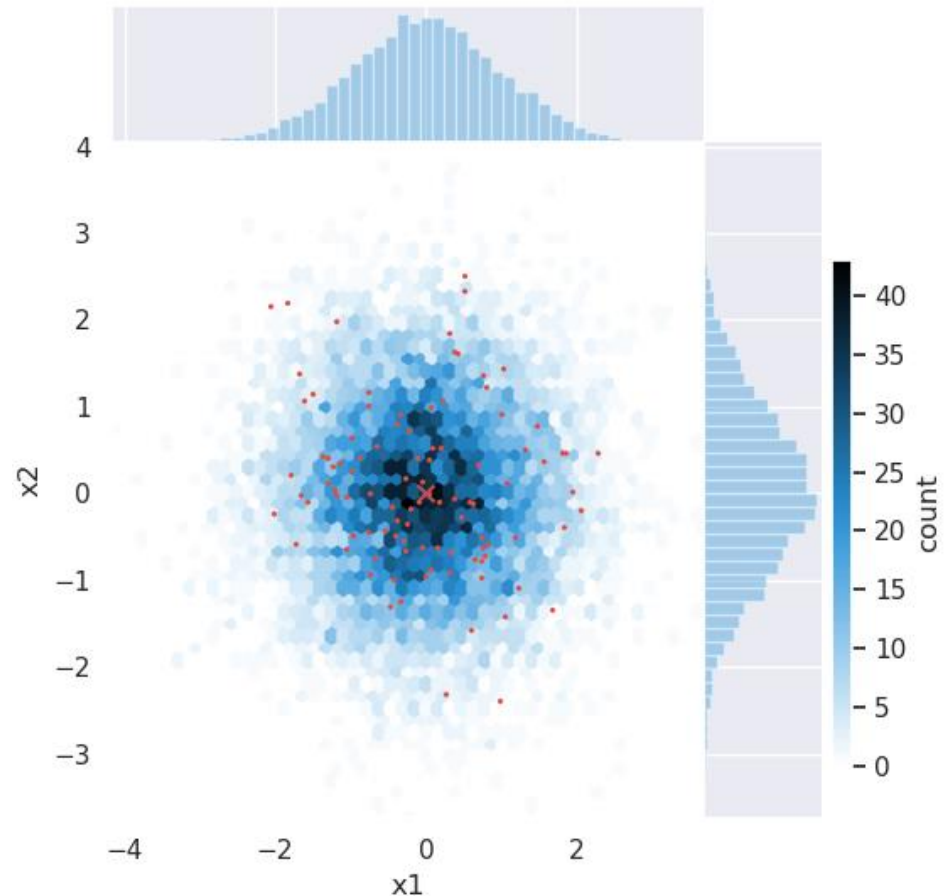
11 μm



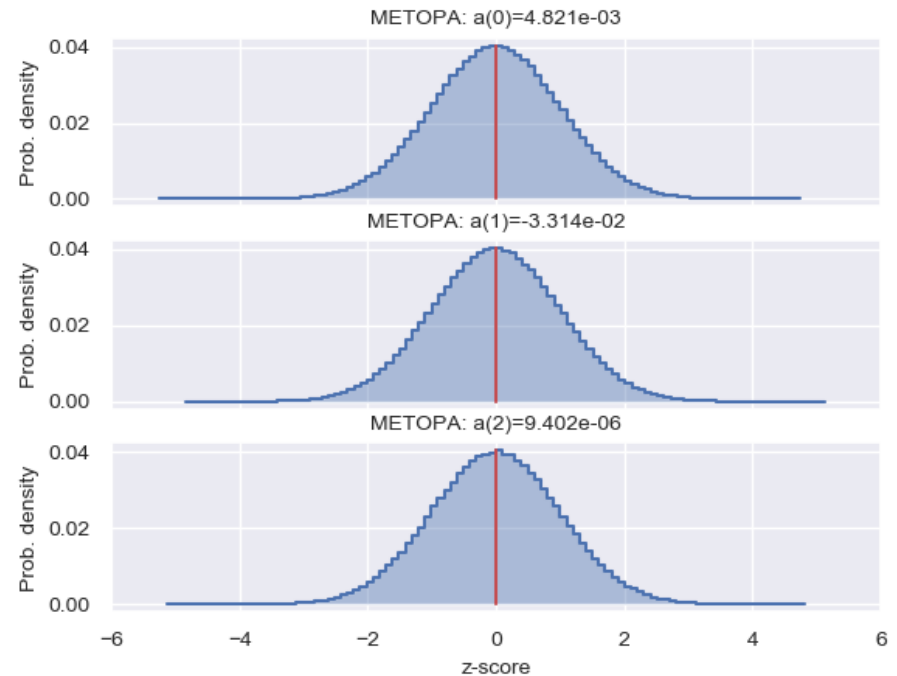
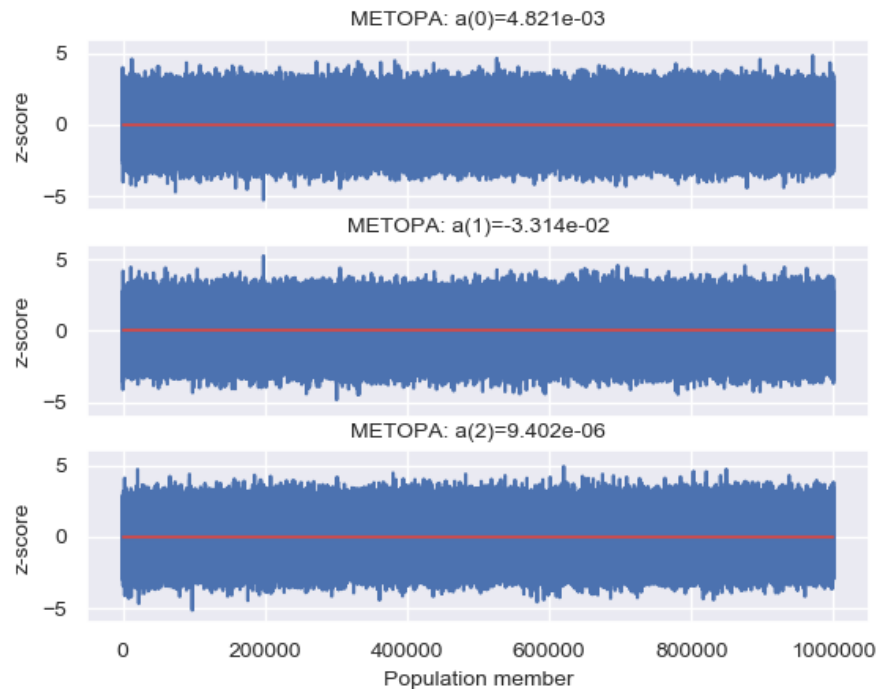
MONTE CARLO: 2D CHECK

Blue = 10000 draws from
bi-normal $X \sim N(\mu=[0,0], \sigma=[1,1])$

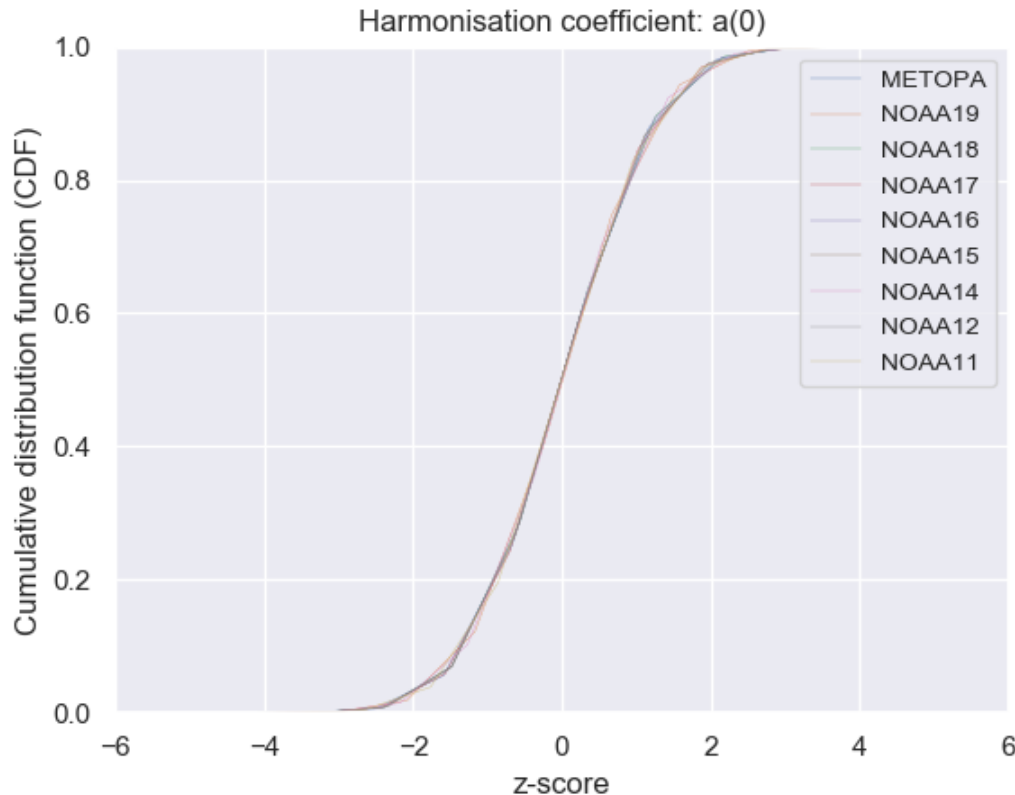
Red = 100 random draws with
`numpy.random.multinormal_normal(mean(X), cov(X), 100)`



Monte Carlo: 27D (& 36D)



METHOD #1: CDF NORM

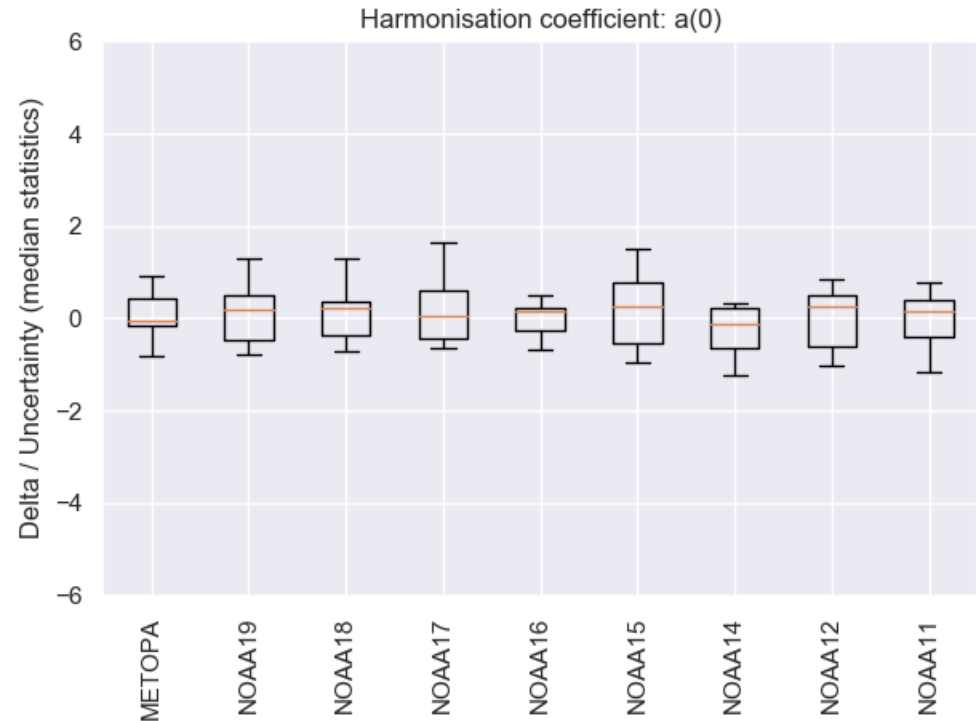
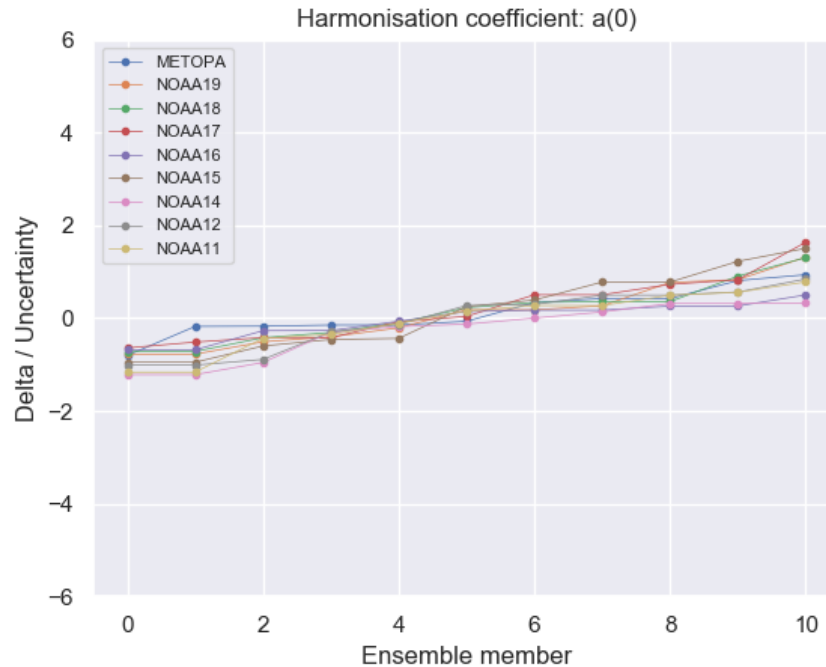


Use sorted draw CDFs
to find deciles
(& store indices)

Loop over all
draw vectors &
calculate norm
for each decile

Index [min(Norm)] → ensemble

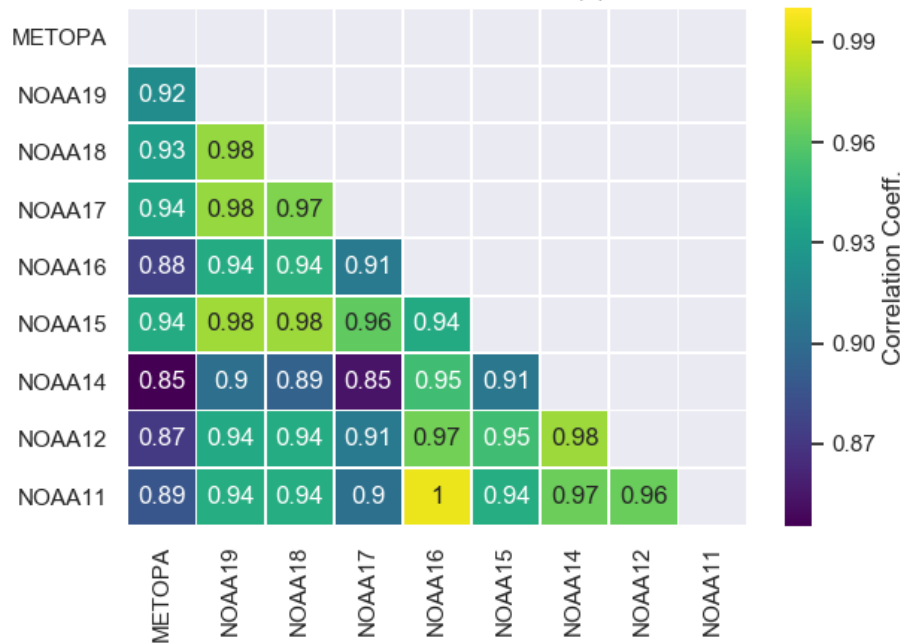
ENSEMBLE STATS



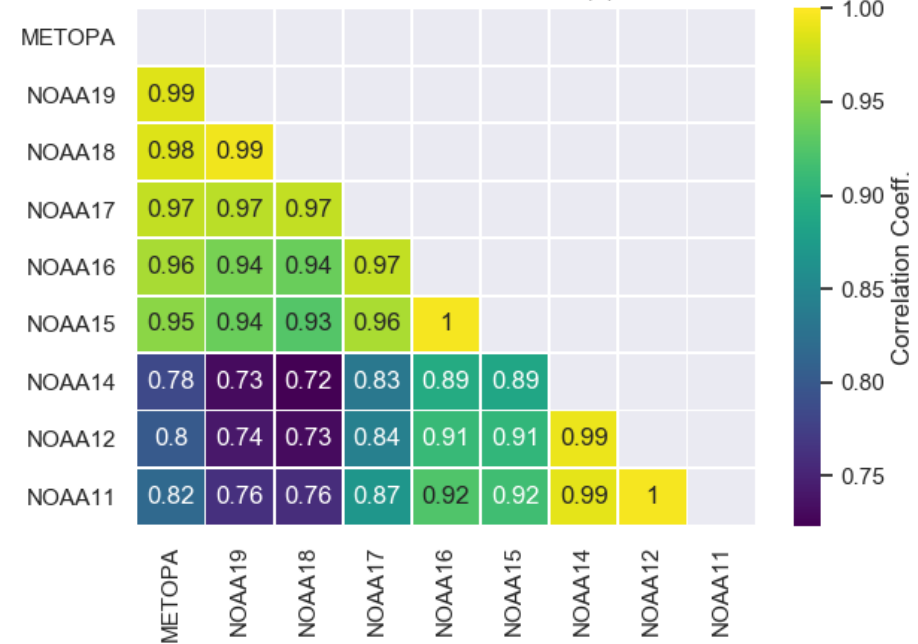
We would like the sorted ensemble to pass through (5,0) and present a CDF-shaped spread – it more or less does

ENSEMBLE CORRELATIONS

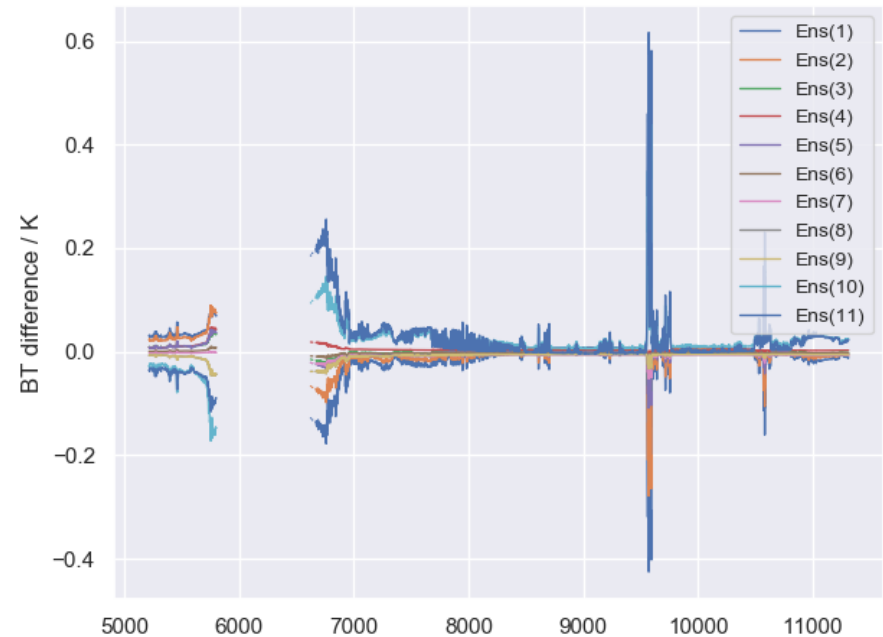
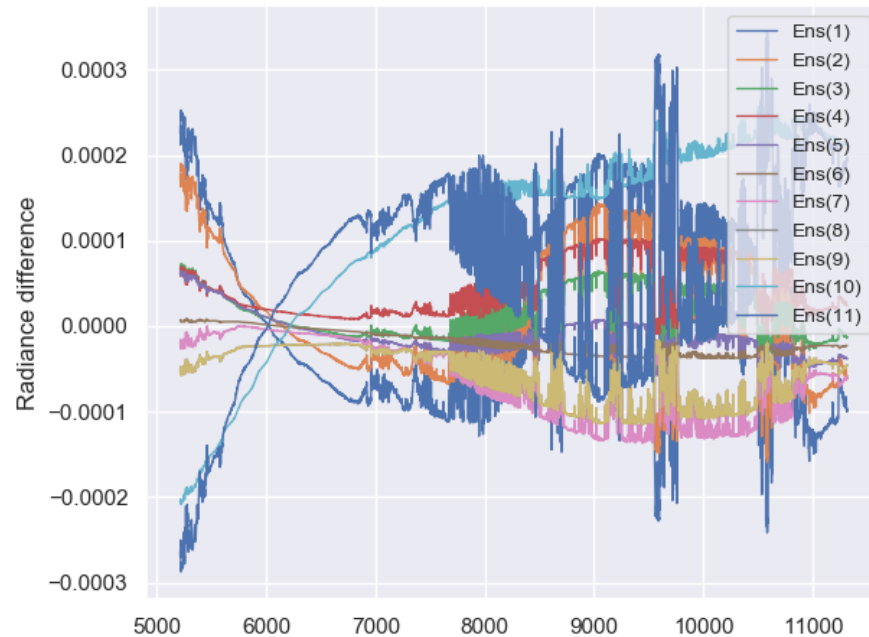
Harmonisation coefficient: $a(0)$



Harmonisation coefficient: $a(1)$



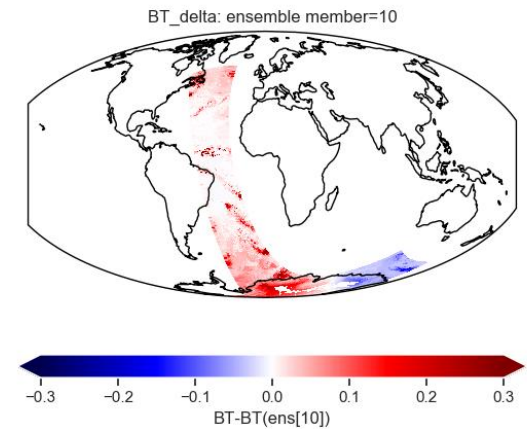
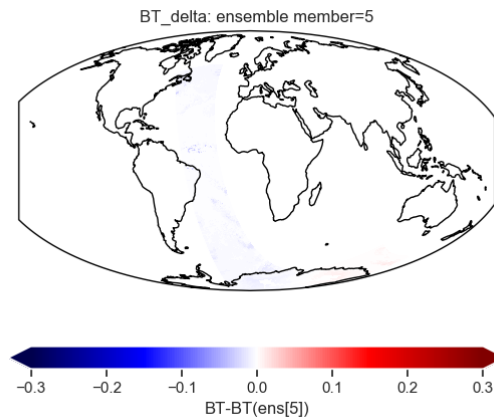
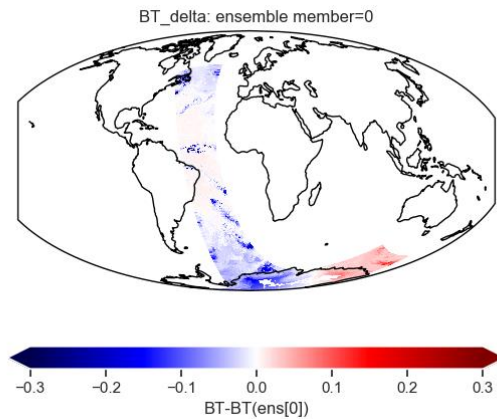
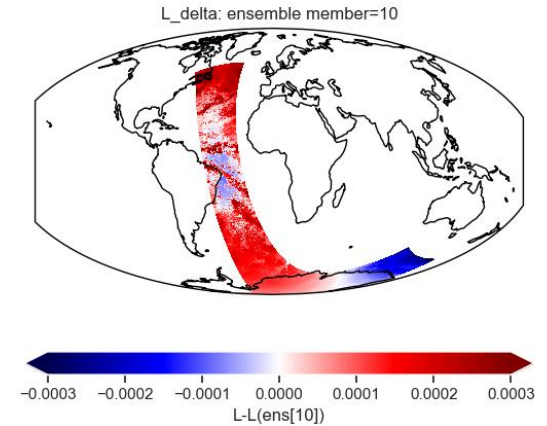
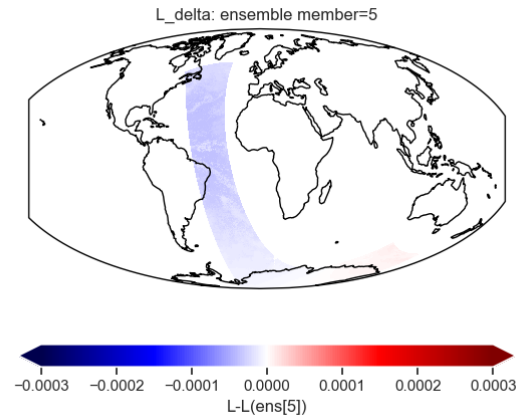
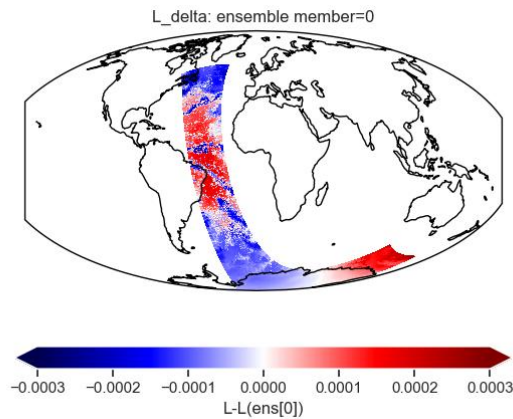
ENSEMBLE RADIANCE



Using L1B counts and temperatures from an orbit from MetOp-A, I calculated the radiance for each ensemble member and then used LUTs from the L1C Easy FCDR (thanks James!) to convert to BTs:

https://github.com/FIDUCEO/MMD_HARM

ORBITAL ENSEMBLE DELTAS



METHOD #2: PCA-MC CHECK

Suppose there are n-samples of p-variables $\rightarrow [n \times p]$ matrix

$$X = \begin{bmatrix} x_{11} & x_{12} & \cdots & x_{1p} \\ x_{21} & x_{22} & \cdots & x_{2p} \\ \vdots & \vdots & & \vdots \\ x_{n1} & x_{n2} & \cdots & x_{np} \end{bmatrix}$$

PCA

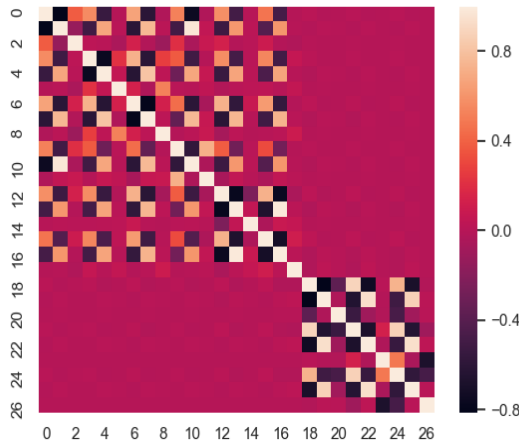


$$\begin{cases} z_1 = l_{11}x_1 + l_{12}x_2 + \cdots + l_{1p}x_p \\ z_2 = l_{21}x_1 + l_{22}x_2 + \cdots + l_{2p}x_p \\ \dots\dots\dots \\ z_m = l_{m1}x_1 + l_{m2}x_2 + \cdots + l_{mp}x_p \\ l_{i1}^2 + \cdots + l_{ip}^2 = 1 \end{cases}$$

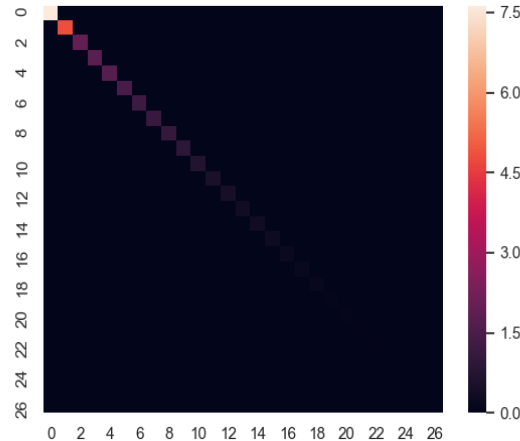
z_m = PCs

x_p = vectors in X

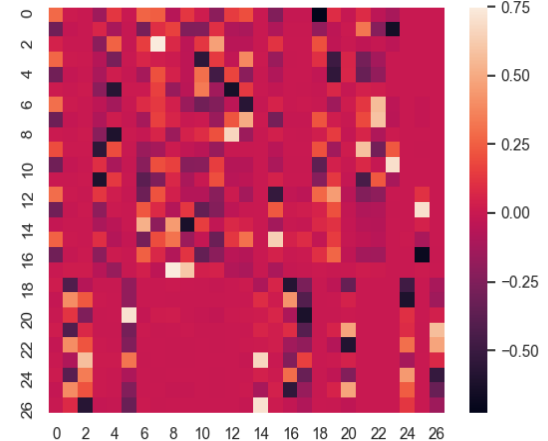
Correlation matrix, R



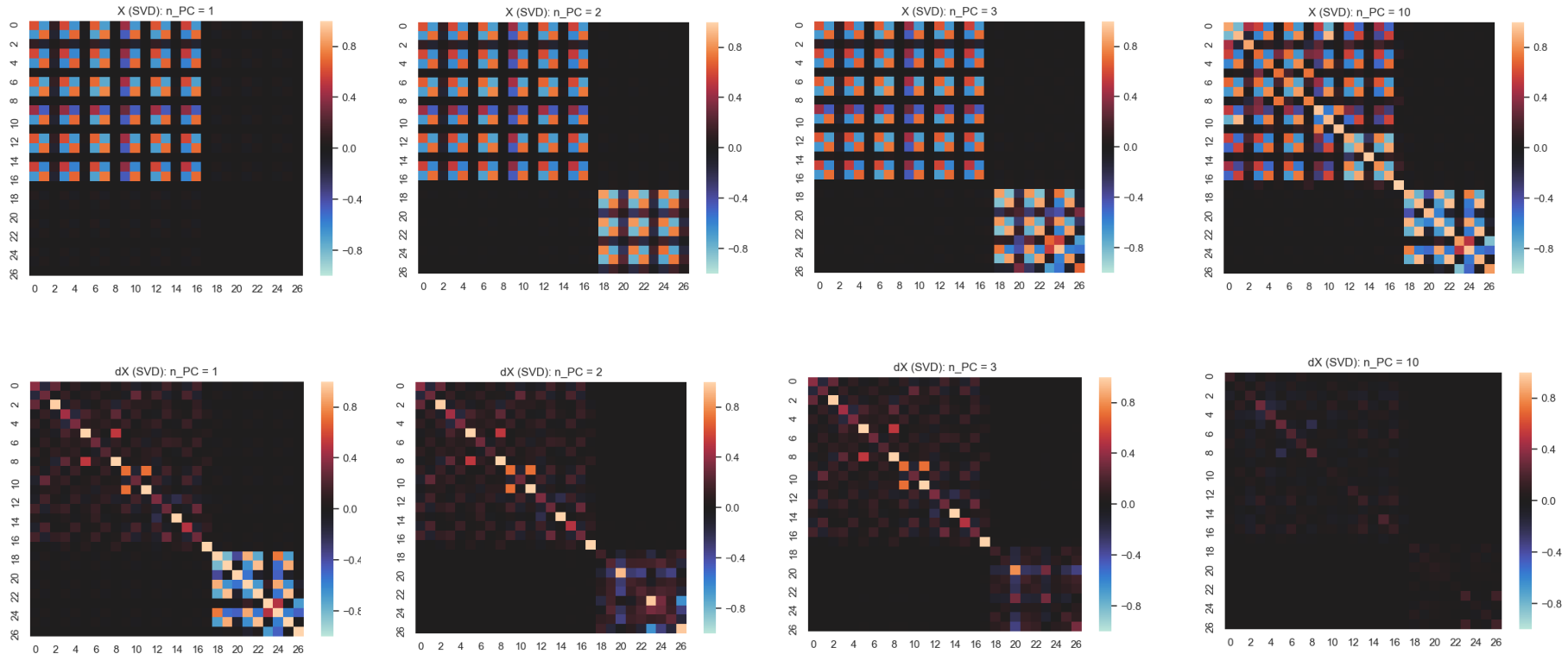
Singular value matrix, S



Eigenvector matrix, E



PCA



Q: Can we sample from the eigenvectors to generate the ensemble directly (rather than looping over all draws and then using the norm with respect to CDF deciles)?

THANKS!

I've put the python code here for re-use / revision:
https://github.com/patternizer/ENSEMBLE_SST